

Pre-geometric Gravity

Emergent Gravity from Topological Quantum Field Theory: Stochastic Gradient Flow Perspective away from the Quantum Gravity Problem

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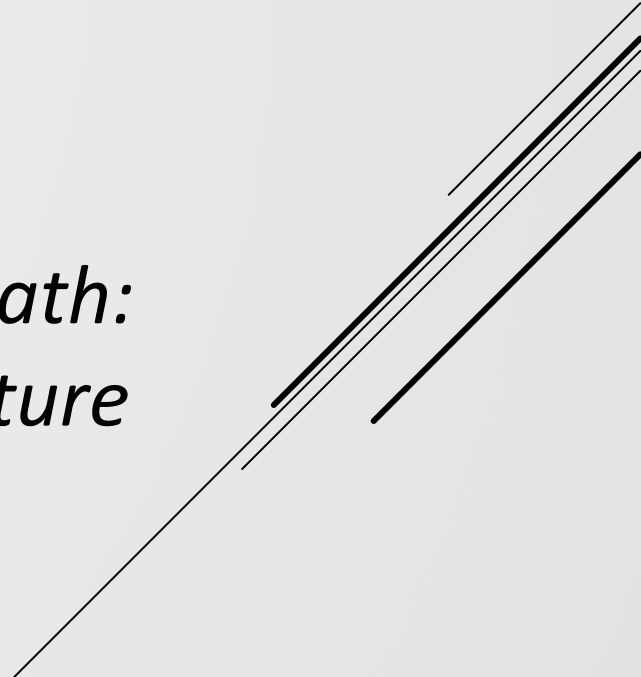
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The problem of Quantum Gravity

*An alternative approach off the beaten path:
gauge theory and emergent metric structure*

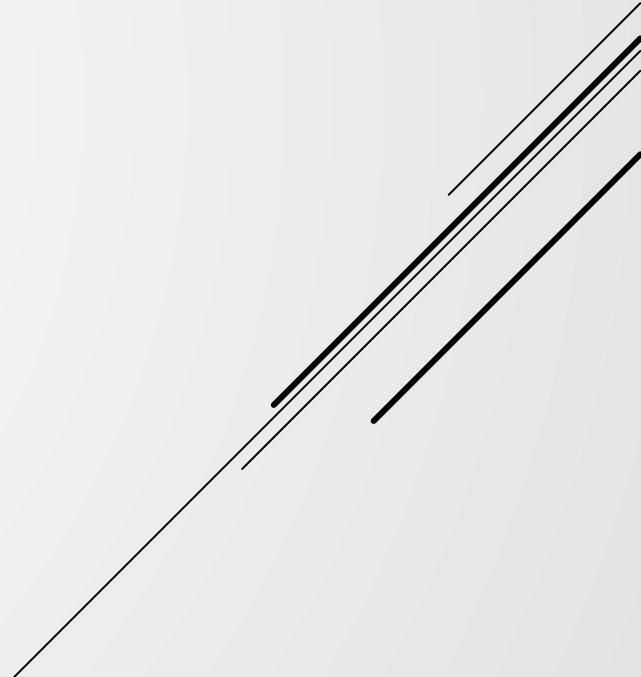


References

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- ▶ Addazi *et al.* (2025) [arXiv:2505.01272](https://arxiv.org/abs/2505.01272)
- ▶ Addazi *et al.* (2025), [arXiv:2505.17014](https://arxiv.org/abs/2505.17014)
- ▶ MacDowell & Mansouri (1977) [10.1103/PhysRevLett.38.739](https://doi.org/10.1103/PhysRevLett.38.739)
- ▶ Pagels (1984) [10.1103/PhysRevD.29.1690](https://doi.org/10.1103/PhysRevD.29.1690)
- ▶ Wilczek (1998) [10.1103/PhysRevLett.80.4851](https://doi.org/10.1103/PhysRevLett.80.4851)
- ▶ Wise (2010) [10.1088/0264-9381/27/15/155010](https://doi.org/10.1088/0264-9381/27/15/155010)
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- ▶ Westman & Zlosnik (2013) [10.1016/j.aop.2013.03.012](https://doi.org/10.1016/j.aop.2013.03.012)

Outline

- ▶ physical motivations
- ▶ constructing the theory
- ▶ Hamiltonian analysis
- ▶ Topological BF theory and Gradient Flow
- ▶ conclusions and perspectives



Quantum Mechanics

$$\phi, \quad \psi^a, \quad A_\mu^a$$

$$[\hat{a}, \hat{a}^\dagger] = \mathbb{I}$$

$$S = \mathbb{I} + iT$$

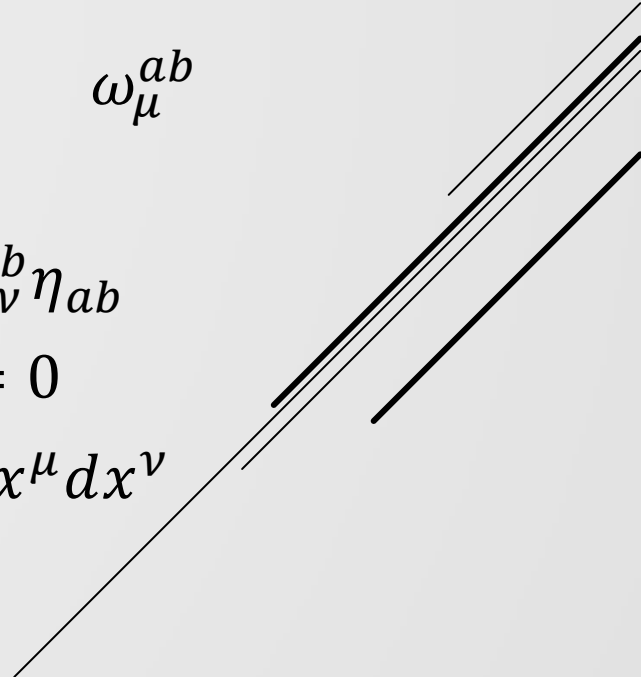
$$\frac{d}{dt} \langle \hat{O} \rangle = \frac{1}{i\hbar} \langle \hat{O}, \hat{H} \rangle + \left\langle \frac{\partial}{\partial t} \hat{O} \right\rangle$$

General Relativity

$$g_{\mu\nu}, \quad e_\mu^a, \quad \omega_\mu^{ab}$$

$$g_{\mu\nu} = e_\mu^a e_\nu^b \eta_{ab}$$

$$\nabla_\lambda g_{\mu\nu} = 0$$

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$
A series of parallel diagonal lines extending from the bottom right towards the center of the slide, serving as a decorative element.

Quantum Mechanics

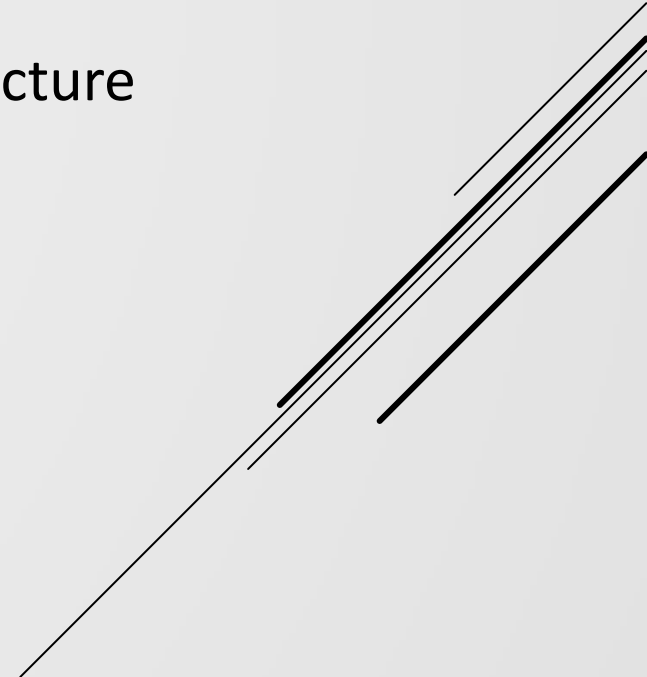
General Relativity

gauge theory

metric structure



spontaneous symmetry breaking



local Lorentz transformation $\Lambda_b^a(x)$

Spin Connections

$$\omega_\mu^{ab} \rightarrow \Lambda_c^a \omega_\mu^{cd} (\Lambda^{-1})_d^b + \eta^{cd} \Lambda_c^a \partial_\mu (\Lambda^{-1})_d^b$$

tetrads

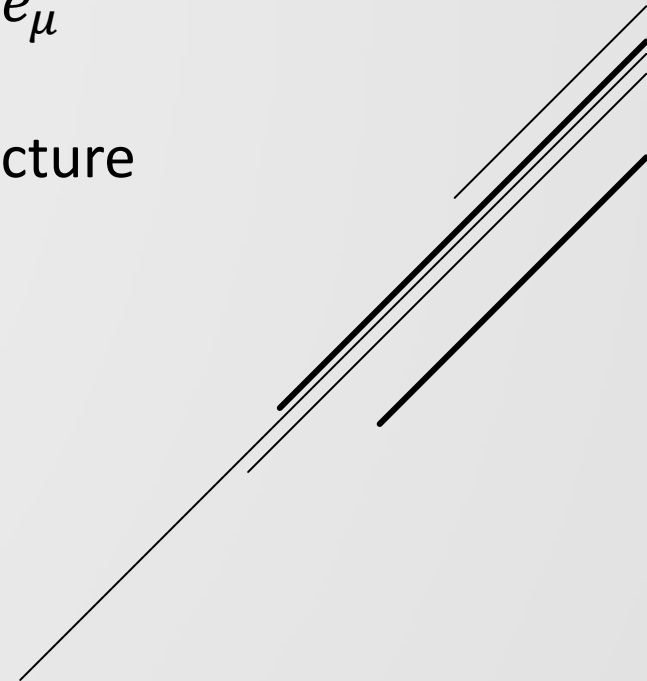
$$e_\mu^a \rightarrow \Lambda_b^a e_\mu^b$$

gauge theory

metric structure



gravitationalise the quantum



Unification Logic

$$40 \ (4 \mu \cdot 10 \ A,B)$$

gauge field A_{μ}^{AB}

$$\omega_{\mu}^{ab}$$

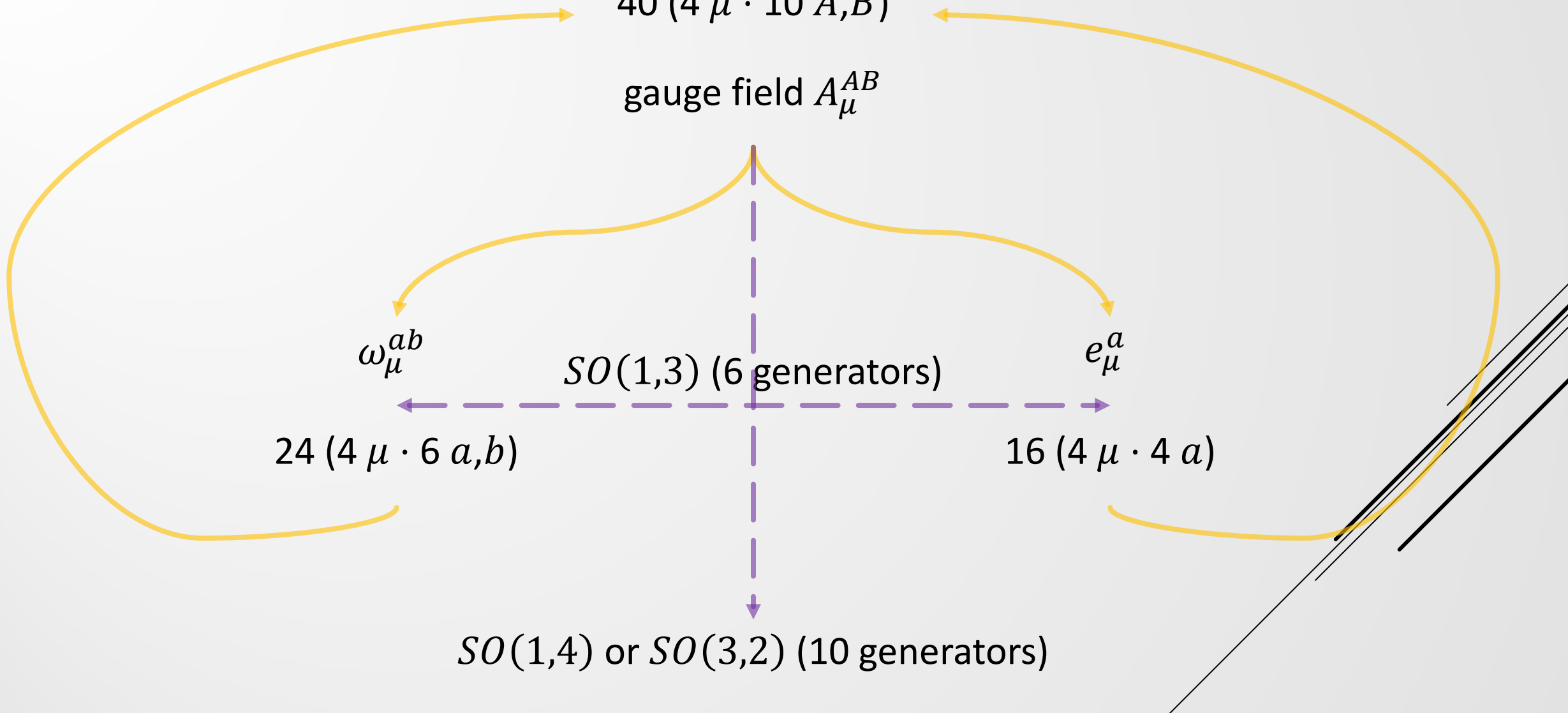
$$24 \ (4 \mu \cdot 6 \ a,b)$$

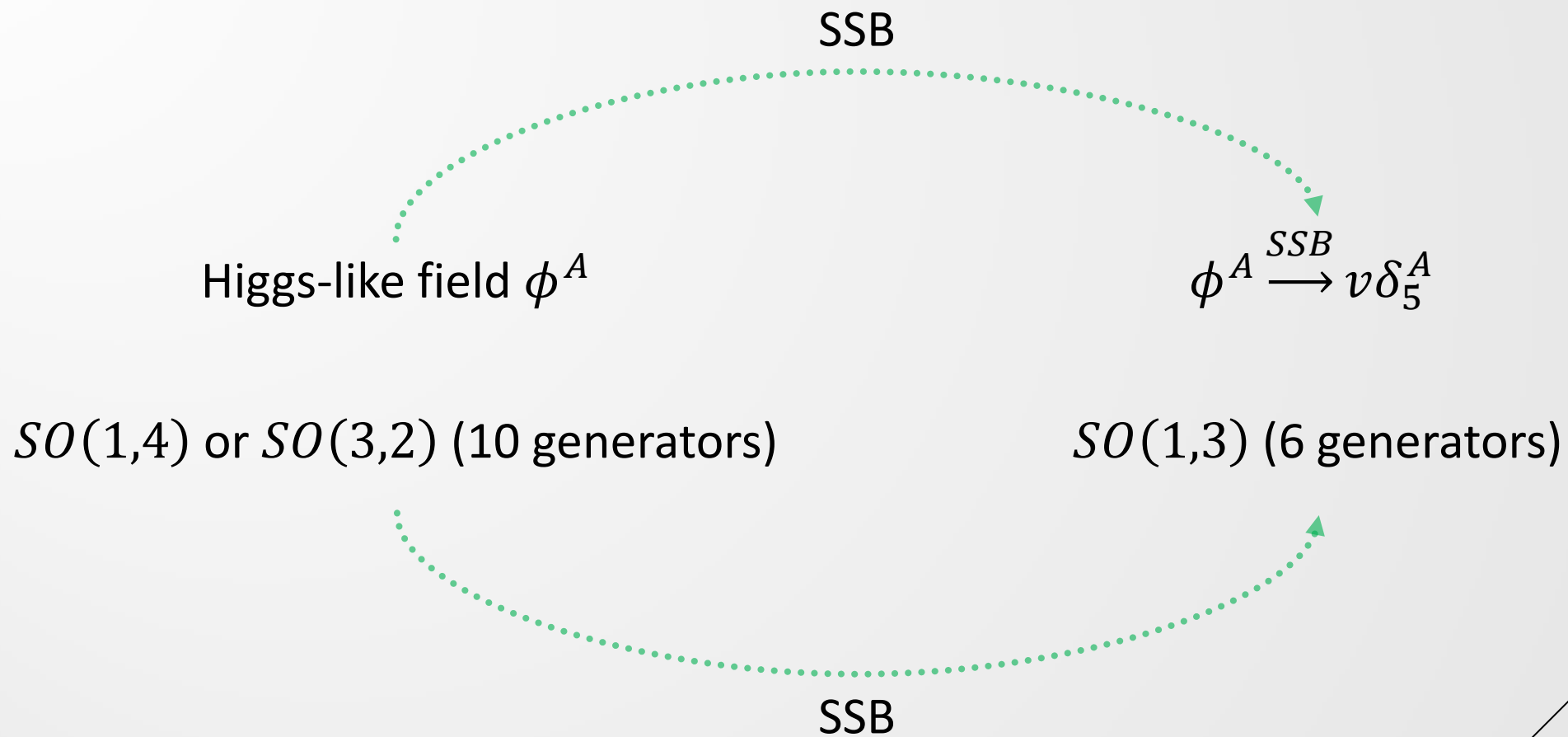
$SO(1,3)$ (6 generators)

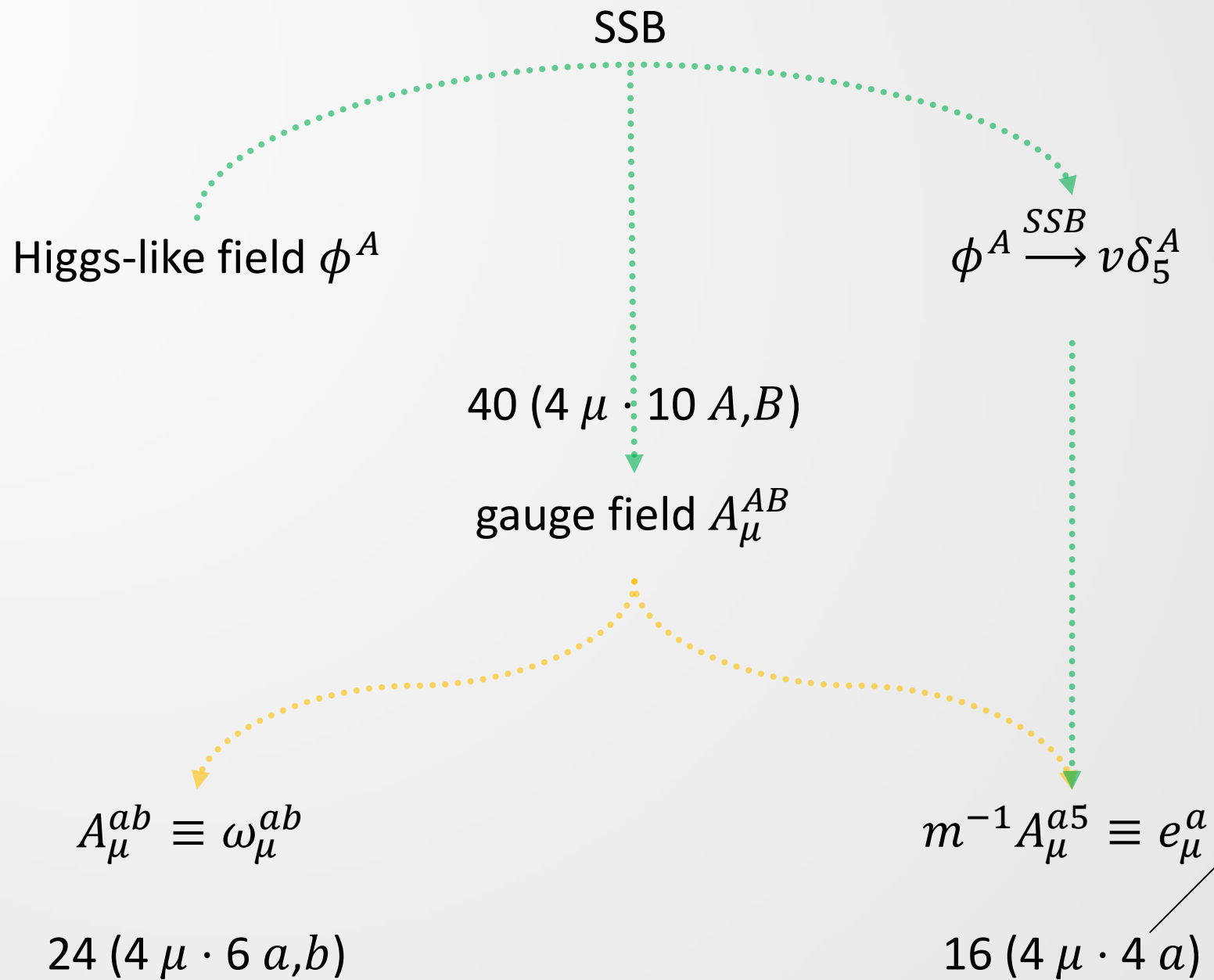
$$e_{\mu}^a$$

$$16 \ (4 \mu \cdot 4 \ a)$$

$SO(1,4)$ or $SO(3,2)$ (10 generators)







$$A_\mu^{AB}, \quad \phi^A$$

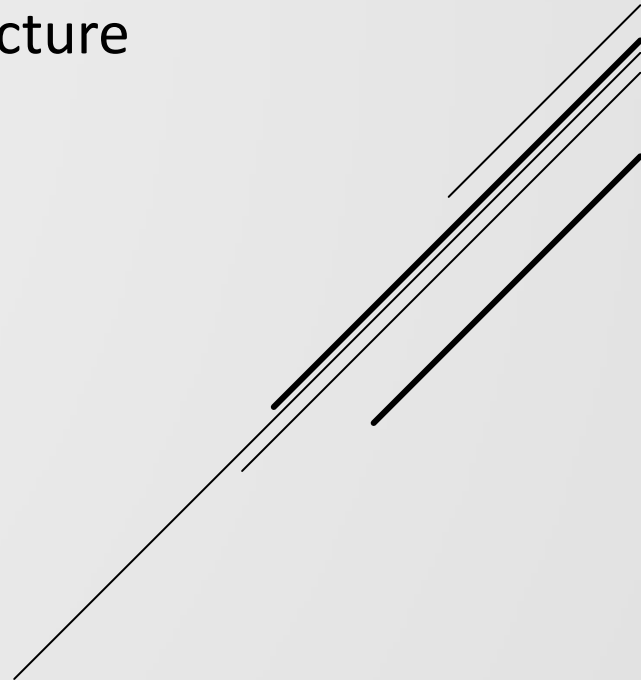
gauge theory

$$g_{\mu\nu}, \quad e_\mu^a, \quad \omega_\mu^{ab}$$

metric structure

SSB

gravitationalise the quantum



SSB

internal-space generalised
Minkowski metric η_{AB}

tangent-space
Minkowski metric η_{ab}

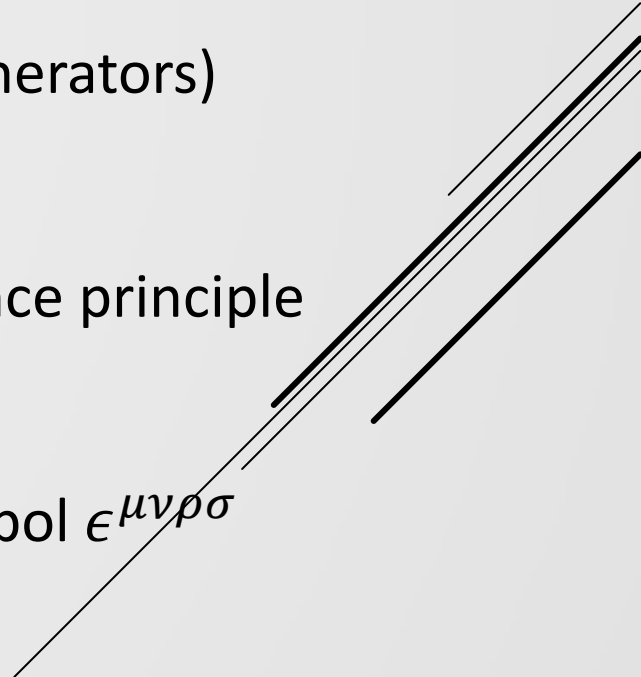
$SO(1,4)$ or $SO(3,2)$ (10 generators)

$SO(1,3)$ (6 generators)

Einstein equivalence principle

no metric structure ~~$g_{\mu\nu}$~~

Levi-Civita symbol $\epsilon^{\mu\nu\rho\sigma}$



Assumptions

- ▶ 4D spacetime
- ▶ principle of general covariance
 - ▶ $\mathcal{L}_{\text{EH}}$ and EEP after SSB
 - ▶ only A_{μ}^{AB} and ϕ^A before SSB
- ▶ polynomiality in A_{μ}^{AB} and ϕ^A and their derivatives

Classical gravity

$$\mathcal{L}_{\text{EH}} = \frac{M_P^2}{2} e e_a^\mu e_b^\nu R_{\mu\nu}^{ab}$$

$R_{\mu\nu}^{ab} = 2(\partial_{[\mu} \omega_{\nu]}^{ab} + \omega_{c[\mu}^a \omega_{\nu]}^{cb})$

$g_{\mu\nu} = e_\mu^a e_\nu^b \eta_{ab}$

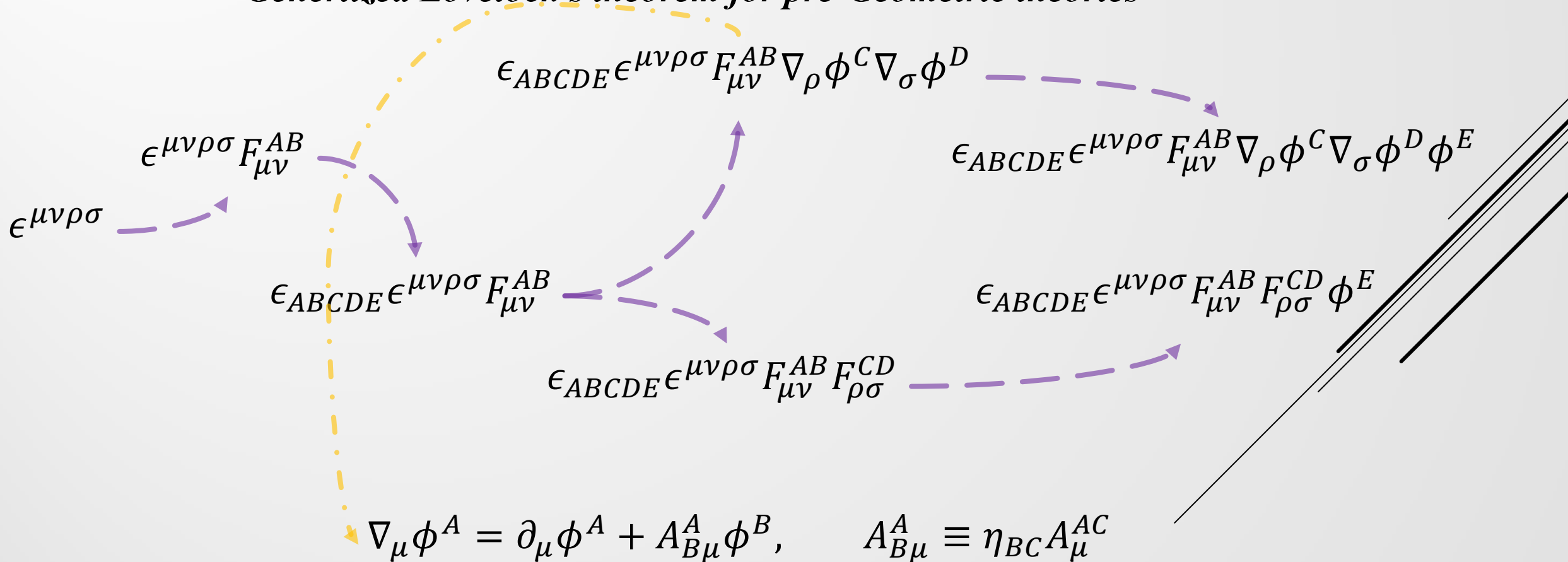
$e \equiv \det e_\mu^a$

$$\mathcal{G} = R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} = (e_a^\mu e_b^\nu e_c^\rho e_d^\sigma - 4e_a^\mu e_d^\nu e_c^\rho e_b^\sigma + e_c^\mu e_d^\nu e_a^\rho e_b^\sigma) R_{\mu\nu}^{ab} R_{\rho\sigma}^{cd}$$

Constructing the theory

rigidity, uniqueness, no so much space for arbitrary lagrangians (Addazi et al 2025)

Generalized Lovelock's theorem for pre-Geometric theories



Two possible Lagrangians

$$\mathcal{L}_W = k_W \epsilon_{ABCDE} \epsilon^{\mu\nu\rho\sigma} F_{\mu\nu}^{AB} \nabla_\rho \phi^C \nabla_\sigma \phi^D \phi^E$$

$$\mathcal{L}_{MM} = k_{MM} \epsilon_{ABCDE} \epsilon^{\mu\nu\rho\sigma} F_{\mu\nu}^{AB} F_{\rho\sigma}^{CD} \phi^E$$

⋮
↓

$$\mathcal{L}_W \xrightarrow{SSB} \mathcal{L}_{EH} - M_P^2 \Lambda e$$

⋮
↓

$$\mathcal{L}_{MM} \xrightarrow{SSB} \mathcal{L}_{EH} - M_P^2 \Lambda e + \lambda e \mathcal{G}$$

↪

$$M_P^2 \equiv -8k_W v^3 m^2$$

$$\Lambda \equiv \pm 6m^2 = \mp \frac{3M_P^2}{4k_W v^3}$$

↪

$$k_W \sim -1 [\phi]^{-3} \Rightarrow v \sim 10^{40} [\phi]$$

↪

$$M_P^2 \equiv \pm 32k_{MM} v m^2$$

$$\Lambda \equiv \pm 3m^2 = \frac{3M_P^2}{32k_{MM} v}$$

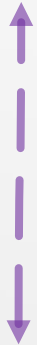
$$\lambda \equiv -4k_{MM} v$$

↪

$$k_{MM} \sim \pm 1 [\phi]^{-1} \Rightarrow v \sim 10^{119} [\phi]$$

Emergence of (non)equivalent principles

General covariance



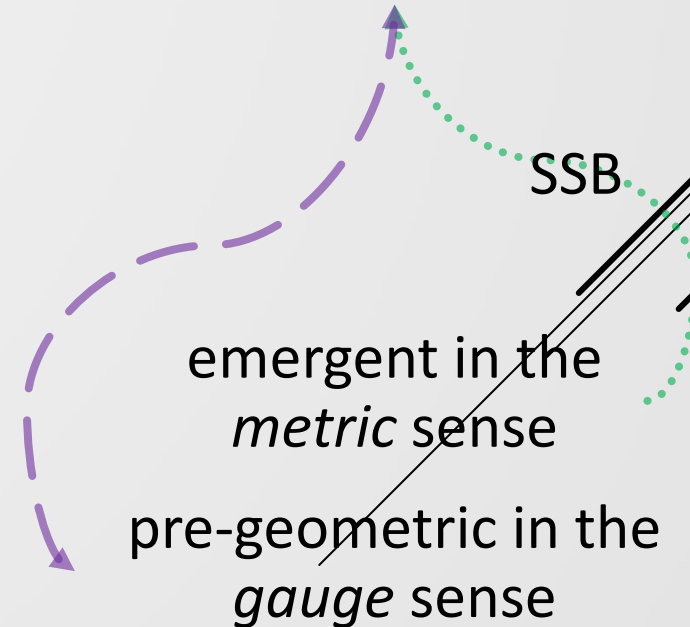
pre-geometric

Diffeomorphism invariance



emergent

Background independence



A dictionary for emergent gravity

effective geometric quantities

$$\pm v^{-1} m^{-1} \nabla_{\mu} \phi^A \xrightarrow{SSB} e_{\mu}^a$$

$$4vmJ^{-1} w_A^{\mu} \xrightarrow{SSB} e_a^{\mu}$$

$$-24^{-1} v^{-5} m^{-4} J \xrightarrow{SSB} e = \sqrt{-g}$$

$$v^{-2} m^{-2} P_{\mu\nu} \xrightarrow{SSB} \eta_{ab} e_{\mu}^a e_{\nu}^b \equiv g_{\mu\nu}$$

$$16v^2 m^2 J^{-2} \eta^{AB} w_A^{\mu} w_B^{\nu} \xrightarrow{SSB} \eta^{ab} e_a^{\mu} e_b^{\nu} \equiv g^{\mu\nu}$$

auxiliary fields

$$P_{\mu\nu} \equiv \eta_{AB} \nabla_{\mu} \phi^A \nabla_{\nu} \phi^B$$

$$w_A^{\mu} \equiv \pm \epsilon_{ABCDE} \epsilon^{\mu\nu\rho\sigma} \nabla_{\nu} \phi^B \nabla_{\rho} \phi^C \nabla_{\sigma} \phi^D \phi^E$$

$$J \equiv \epsilon_{ABCDE} \epsilon^{\mu\nu\rho\sigma} \nabla_{\mu} \phi^A \nabla_{\nu} \phi^B \nabla_{\rho} \phi^C \nabla_{\sigma} \phi^D \phi^E$$

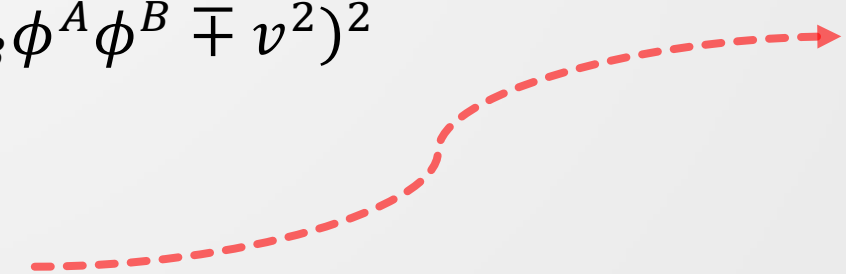
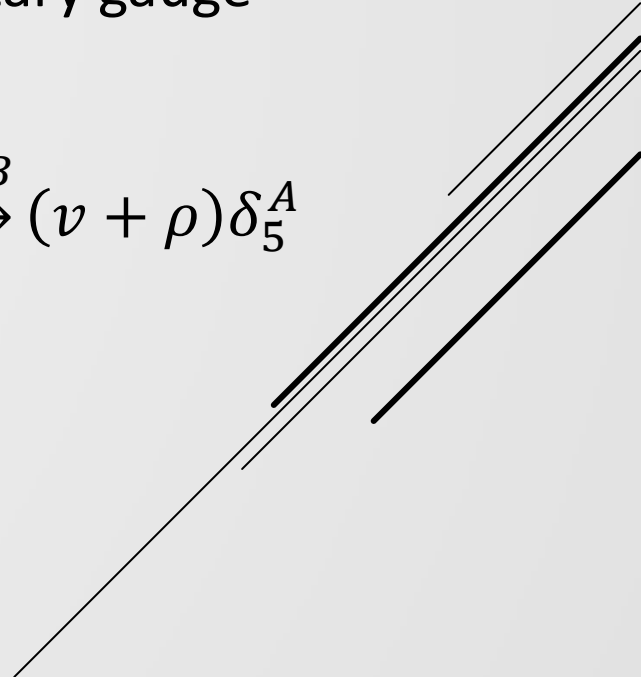
Spontaneous symmetry breaking

symmetry-breaking potential

unitary gauge

$$\mathcal{L}_{\text{SSB}} = -k_{\text{SSB}} v^{-4} |J| (\eta_{AB} \phi^A \phi^B \mp v^2)^2$$


$$\phi^A = v \delta_5^A$$


$$\phi^A \xrightarrow{\text{SSB}} (v + \rho) \delta_5^A$$


Quantum fluctuations of the background geometry

total effective geometric quantities

'classical' and 'quantum' parts

$$\pm v^{-1} m^{-1} \nabla_{\mu} \phi^A \xrightarrow{SSB} \mathcal{E}_{\mu}^a$$

$$4vmJ^{-1} w_A^{\mu} \xrightarrow{SSB} \mathcal{E}_a^{\mu}$$

$$-24^{-1} v^{-5} m^{-4} J \xrightarrow{SSB} \mathcal{E} = \sqrt{-\mathcal{G}}$$

$$v^{-2} m^{-2} P_{\mu\nu} \xrightarrow{SSB} \eta_{ab} \mathcal{E}_{\mu}^a \mathcal{E}_{\nu}^b \equiv \mathcal{G}_{\mu\nu}$$

$$16v^2 m^2 J^{-2} \eta^{AB} w_A^{\mu} w_B^{\nu} \xrightarrow{SSB} \eta^{ab} \mathcal{E}_a^{\mu} \mathcal{E}_b^{\nu} \equiv \mathcal{G}^{\mu\nu}$$

$$\mathcal{E}_{\mu}^a \equiv e_{\mu}^a + \hat{e}_{\mu}^a, \quad \mathcal{E}_a^{\mu} \equiv e_a^{\mu} + \hat{e}_a^{\mu}, \quad \mathcal{E} \equiv e + \hat{e},$$

$$\mathcal{G}_{\mu\nu} \equiv g_{\mu\nu} + \hat{g}_{\mu\nu}, \quad \mathcal{G}^{\mu\nu} \equiv g^{\mu\nu} + \hat{g}^{\mu\nu}$$

quantum fluctuations of the metric

$$\hat{g}_{\mu\nu} = \mathcal{G}_{\mu\nu} - g_{\mu\nu} \xleftarrow{SSB} v^{-2} m^{-2} P_{\mu\nu} - m^{-2} \eta_{ab} A_{\mu}^{a5} A_{\nu}^{b5}$$

Matter couplings and the equivalence principle

correspondence principle

$$\mathcal{L}_\Phi = \frac{1}{3} v^{-3} m^{-2} J^{-1} w_A^\mu w_B^\nu \eta^{AB} \partial_\mu \Phi \partial_\nu \Phi \xrightarrow{SSB} -\frac{1}{2} \sqrt{-g} g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi$$

$$\mathcal{L}_\Psi = -\frac{i}{6} v^{-4} m^{-3} \bar{\Psi} w_A^\mu \gamma^A \nabla_\mu \Psi \xrightarrow{SSB} i \sqrt{-g} \bar{\Psi} \gamma^\mu \nabla_\mu \Psi$$

$$\mathcal{L}_G = \frac{8}{3} v^{-1} J^{-3} w_A^\rho w_B^\mu w_C^\sigma w_D^\nu \eta^{AB} \eta^{CD} \text{tr}(G_{\mu\nu} G_{\rho\sigma}) \xrightarrow{SSB} -\frac{1}{4} \sqrt{-g} g^{\mu\rho} g^{\nu\sigma} \text{tr}(G_{\mu\nu} G_{\rho\sigma})$$

Gravitational Higgs mechanism

kinematics of the Higgs-like field

$$\mathcal{L}_\phi = \frac{1}{3} v^{-3} J^{-1} w_A^\mu w_B^\nu \eta^{AB} \eta_{CD} \nabla_\mu \phi^C \nabla_\nu \phi^D \xrightarrow{SSB} -\frac{1}{2} m^2 \sqrt{-g} g^{\mu\nu} \eta_{AB} \nabla_\mu \phi^A \nabla_\nu \phi^B + \dots$$

$$A_\mu^{ab} \xrightarrow{SSB} \omega_\mu^{ab} \dots \rightarrow \text{massless}$$

$$m^{-1} A_\mu^{a5} \xrightarrow{SSB} e_\mu^a \dots \rightarrow \text{massive, but actually...}$$

$$\phi^A \xrightarrow{SSB} (v + \rho) \delta_5^A \dots \rightarrow \text{massive}$$

Gravitational Higgs mechanism

Λ sector

$$\mathcal{L}_W: \cdots \rightarrow \mathcal{L}_\Lambda \equiv -M_{\text{P}}^2 \left(\pm 6m^2 - \frac{m^2}{4k_W v} \right) e$$

$$\cancel{\mathcal{L}_{\text{MM}}}: \cdots \rightarrow \mathcal{L}_\Lambda \equiv -M_{\text{P}}^2 \left(\pm 3m^2 \pm \frac{vm^2}{16k_{\text{MM}}} \right) e$$

$$\begin{aligned} \mathcal{L}_\phi \xrightarrow{SSB} & -\frac{1}{2} m^2 \sqrt{-g} g^{\mu\nu} \eta_{AB} \nabla_\mu \phi^A \nabla_\nu \phi^B + \cdots = -\frac{1}{2} m^2 \sqrt{-g} g^{\mu\nu} \eta_{ab} (v m e_\mu^a) (v m e_\nu^b) + \cdots \\ & = -\frac{1}{2} m^2 \sqrt{-g} g^{\mu\nu} \eta_{ab} (v^2 m^2) e_\mu^a e_\nu^b + \cdots = -\frac{1}{2} v^2 m^4 \sqrt{-g} g^{\mu\nu} g_{\mu\nu} + \cdots = -2v^2 m^4 \sqrt{-g} + \cdots \end{aligned}$$

Gravitational Higgs mechanism

$SO(1,4)$ ~~$SO(3,2)$~~

ρ sector

$$\mathcal{L}_\rho \equiv \mp \frac{1}{2} m^2 \sqrt{-g} g^{\mu\nu} \partial_\mu \rho \partial_\nu \rho + \sqrt{-g} [12k_W v^2 m^2 (-e_a^\mu e_b^\nu R_{\mu\nu}^{ab} \pm 12m^2) - 4vm^4] \rho$$

$$+ \sqrt{-g} [12k_W v m^2 (-e_a^\mu e_b^\nu R_{\mu\nu}^{ab} \pm 12m^2) - 96k_{SSB} v^3 m^4 - 2m^4] \rho^2 + \mathcal{O}(\rho^3)$$

$$m_\rho^2 = -24 \frac{k_{SSB}}{k_W} M_P^2 + \frac{36}{v^2} M_P^2 + \frac{2}{3} \Lambda \approx (1.43 \cdot 10^{38} k_{SSB} + 2.07 \cdot 10^{-42} + 2.83 \cdot 10^{-84}) \text{ GeV}^2$$

$$m_\rho \sim M_P \text{ for } k_{SSB} \sim 10^{-2} \text{ or } k_{SSB} \ll 10^{116} \text{ and } v \sim 10$$

Cosmological implications

- ▶ eternal and singularity-free pre-geometric universe before SSB
 - ▶ problem of causality
- ▶ phase transition at $T_c \sim m_\rho \lesssim M_P$ if $k_{\text{SSB}} \lesssim 10^{-2}$
 - ▶ supermassive boson ρ

Hamiltonian formulation

conjugate momenta

$$\Pi_A = \frac{\delta(\mathcal{L}_W + \mathcal{L}_{SSB})}{\delta \dot{\phi}^A} = 2\epsilon_{ABCDE}\epsilon^{0ijk}\nabla_k\phi^D\phi^E \left[k_W F_{ij}^{BC} - \frac{2k_{SSB}}{v^4|J|} J \nabla_i\phi^B\nabla_j\phi^C (\eta_{FG}\phi^F\phi^G \mp v^2)^2 \right],$$
$$\Pi_{AB}^i = \frac{\delta(\cancel{\mathcal{L}_W} + \cancel{\mathcal{L}_{SSB}})}{\delta \dot{A}_i^{AB}} = 2k_W\epsilon_{ABCDE}\epsilon^{0ijk}\nabla_j\phi^C\nabla_k\phi^D\phi^E, \quad \Pi_{AB}^0 = \frac{\delta(\cancel{\mathcal{L}_W} + \cancel{\mathcal{L}_{SSB}})}{\delta \dot{A}_0^{AB}} = 0$$

Lagrangian density

$$\mathcal{L}_W + \mathcal{L}_{SSB} = \Pi_{AB}^i F_{0i}^{AB} + \Pi_A \nabla_0 \phi^A$$

Hamiltonian density

$$\mathcal{H}_W = \Pi_{AB}^i \dot{A}_i^{AB} + \Pi_A \dot{\phi}^A - \mathcal{L}_W - \mathcal{L}_{SSB} = \Pi_{AB}^i (\partial_i A_0^{AB} - 2A_{C[0}^A A_{i]}^{CB}) - \Pi_A A_{B0}^A \phi^B$$

Hamiltonian

$$H_W = - \int A_0^{AB} (\partial_i \Pi_{AB}^i + 2 \Pi_{BC}^i A_{Ai}^C + \eta_{BC} \Pi_A \phi^C) d^3x + \int \partial_i (\Pi_{AB}^i A_0^{AB}) d^3x$$



$$H_W \xrightarrow{SSB} \frac{M_P^2}{4} \epsilon_{abcd} \epsilon^{0ijk} \left\{ \int [\omega_0^{ab} \partial_i (e_j^c e_k^d) + 2 \omega_{e0}^a \omega_i^{eb} e_j^c e_k^d] d^3x \right. \\ \left. + \int [\mp 4m^2 e_0^a e_i^b e_j^c e_k^d + e_0^a e_i^b R_{jk}^{cd} - \partial_i (\omega_0^{ab} e_j^c e_k^d)] d^3x \right\}$$

$$\Pi_A \xrightarrow{SSB} \Pi_a = \pm 2k_W v^2 m \epsilon_{abcd} \epsilon^{0ijk} e_k^d (R_{ij}^{bc} \mp 2m^2 e_i^b e_j^c),$$

$$\Pi_{AB}^i \xrightarrow{SSB} \Pi_{ab}^i = 2k_W v^3 m^2 \epsilon_{abcd} \epsilon^{0ijk} e_j^c e_k^d, \quad \Pi_{AB}^0 \xrightarrow{SSB} 0$$

$$\Pi_{AB}^i \dot{A}_i^{AB} + \Pi_A \dot{\phi}^A = \Pi_{ab}^i \dot{A}_i^{ab} + 2m \Pi_{a5}^i \dot{A}_i^{a5} + \Pi_A \dot{\phi}^A$$

$$\xrightarrow{SSB} \Pi_{ab}^i \dot{\omega}_i^{ab} + \cancel{2m \Pi_{a5}^i \dot{e}_i^a} + \cancel{\dot{\phi} \Pi_A \delta_5^A}$$

$$\mathcal{L}_W \xrightarrow{SSB} \mathcal{L}_{EH} + \mathcal{L}_\Lambda$$

$$\mathcal{L}_{SSB} \xrightarrow{SSB} 0$$

$$\mathcal{H}_W \xrightarrow{SSB} \Pi_{ab}^i \dot{\omega}_i^{ab} - \mathcal{L}_{EH} - \mathcal{L}_\Lambda$$

ADM formalism

$$\mathcal{H}_{ADM} = \pi_{ab}^i \dot{\omega}_i^{ab} - \mathcal{L}_{EH} - \mathcal{L}_\Lambda$$

$$\Pi_{ab}^i = \pi_{ab}^i \Leftrightarrow \text{time gauge}$$

recovery of the canonical formulation of General Relativity after SSB

recovery of the canonical formulation of General Relativity after SSB

$$\begin{aligned}
 H_W &\xrightarrow{SSB} \frac{M_P^2}{4} \epsilon_{abcd} \epsilon^{0ijk} \left\{ \int \left[\omega_0^{ab} \partial_i (e_j^c e_k^d) + 2 \omega_{e0}^a \omega_i^{eb} e_j^c e_k^d \right] d^3x \right. \\
 &\quad \left. + \int \left[\mp 4m^2 e_0^a e_i^b e_j^c e_k^d + e_0^a e_i^b R_{jk}^{cd} - \partial_i (\omega_0^{ab} e_j^c e_k^d) \right] d^3x \right\} \\
 &\quad \mathcal{H}_\Lambda = M_P^2 \Lambda e \qquad N \mathcal{H}_\perp + N^i \mathcal{H}_i \qquad \partial_i (\pi_{ab}^i \omega_0^{ab}) \\
 &\quad H_{\text{ADM}} = \int \left[N \mathcal{H}_\perp + N^i \mathcal{H}_i + \omega_0^{ab} \mathcal{J}_{ab} + \mathcal{H}_\Lambda + \partial_i (\pi_{ab}^i \omega_0^{ab}) \right] d^3x
 \end{aligned}$$

Loop Quantum Gravity and BF theories

$$\Pi_{A0}^i \xrightarrow{SSB} \Pi_{a0}^i = \frac{M_P^2}{4} \epsilon_{0abc} \epsilon^{0ijk} e_j^b e_k^c = \frac{M_P^2}{2} \tilde{E}_a^i = \frac{M_P^2}{2} \sqrt{h} E_a^i \text{ Ashtekar's "electric field"}$$

$$S_{\text{EH}} = \frac{M_P^2}{2} \int \text{tr}(B \wedge F) = \frac{M_P^2}{4} \int \epsilon_{abcd} e^c \wedge e^d \wedge R^{ab}$$

$$\Leftrightarrow B_{ab}^{(\text{EH})} = \frac{M_P^2}{4} \epsilon_{abcd} e^c \wedge e^d \text{ simplicity constraint}$$

$$H_W \xrightarrow{SSB} H_{\text{ADM}} \Leftrightarrow \text{time gauge}$$

$$S_W = \int \text{tr}(B \wedge F) = -2k_W \int \epsilon_{ABCDE} \nabla \phi^C \wedge \nabla \phi^D \phi^E \wedge F^{AB}$$

$$\Leftrightarrow B_{AB}^{(\text{W})} = -2k_W \epsilon_{ABCDE} \nabla \phi^C \wedge \nabla \phi^D \phi^E \xrightarrow{SSB} B_{ab}^{(\text{W})} = B_{ab}^{(\text{EH})}$$

Algebra of constraints

extended Hamiltonian density

$$\begin{aligned}\mathcal{H}_W^{ext} &= -A_0^{AB} \left(\partial_i \Pi_{AB}^i + 2\Pi_{BC}^i A_{Ai}^C + \eta_{BC} \Pi_A \phi^C \right) + \lambda^A Z_A + \lambda_i^{AB} Z_{AB}^i + \lambda_0^{AB} Z_{AB}^0 \\ &= \lambda^A Z_A + \lambda_i^{AB} Z_{AB}^i + \lambda_0^{AB} Z_{AB}^0 + \tilde{\lambda}_0^{AB} \dot{Z}_{AB}^0\end{aligned}$$

 A_0^{AB} is a *gauge* degree of freedom (λ_0^{AB} is undetermined)

constraint	first-class	second-class
primary	Z_{AB}^0	Z_A, Z_{AB}^i
secondary		$\dot{Z}_{AB}^0$

Degrees of freedom

► 90 dynamical variables: $10 A_0^{AB}$, $30 A_i^{AB}$, $5 \phi^A$, $10 \Pi_{AB}^0$, $30 \Pi_{AB}^i$ and $5 \Pi_A$

► 20 gauge choices: $-10 A_0^{AB}$ and $-10 \Pi_{AB}^0$

► 10 independent first-class constraints: $10 Z_{AB}^0$

► 44 independent second-class constraints: $30 Z_{AB}^i$, $5 Z_A$, $10 \dot{Z}_{AB}^0$ and $-1 H_W^{ext}$

2#degrees of freedom

= #variables – 2#first class constraints – #second class constraints

$$= 90 - 20 - 20 - 44 = 6$$

3 degrees of freedom: $h_{\mu\nu}$ and ρ (like in scalar-tensor *metric* theories)

$$S_{\text{P-W}} = \int (B_{AB} \wedge F^{AB} + \epsilon_{ABCDE} B^{AB} \wedge B^{CD} \phi^E)$$

$$B^{(AB} \wedge B^{CD)} = 0,$$

$$\frac{\partial A^{AB}}{\partial s} = -i\mathcal{D}B^{AB} + \xi_g A^{AB},$$

$$\frac{\partial B^{AB}}{\partial s} = -iF^{AB} - i\epsilon^{AB}{}_{CDE} B^{CD} \phi^E + \xi_f B^{AB},$$

$$\frac{\partial \phi^A}{\partial s} = -i\epsilon^A{}_{BCDE} B^{BC} \wedge B^{DE} + \xi_s \phi^A,$$

$$\frac{\partial \phi^A}{\partial s} - \xi_s \phi^A \approx 0.$$

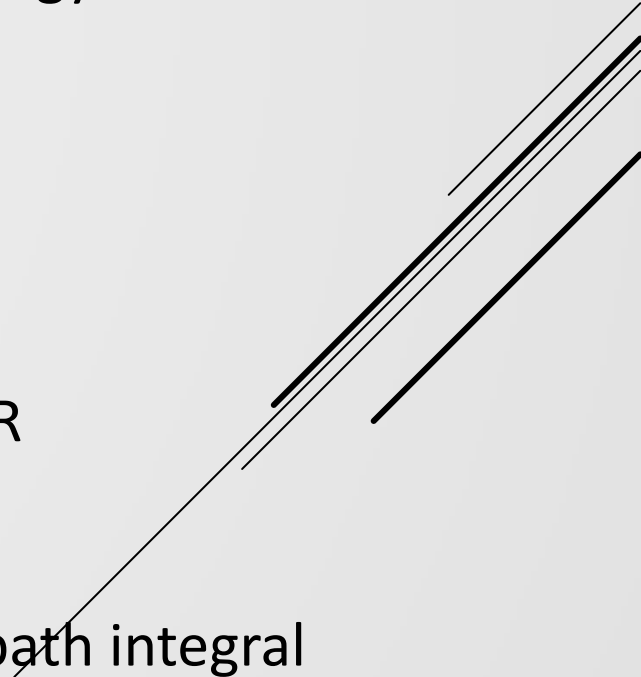
$$\phi^A(s) \approx e^{-\frac{\sigma_{\xi_s}^2}{2}(s-s_0)} \phi_0^A(s),$$

$$\lim_{s \rightarrow s_0} \phi_0^A(s) \sim \frac{v}{2} \delta_5^A [1 + \theta(s - s_0)].$$

Conclusions

- ▶ gravity can emerge from SSB in a pre-geometric universe
- ▶ $SO(1,4) \xrightarrow{SSB} SO(1,3)$: emerging principles and energy scales of GR from $\mathcal{L}_W$
 - ▶ phase transition at $T_c \lesssim M_P$ and supermassive boson ρ
 - ▶ gravity as a BF theory with 2+1 degrees of freedom

Perspectives

- ▶ resolution of classical spacetime singularities, Quantum Cosmology and Inflation
 - ▶ causality, propagators and matter couplings
 - ▶ quantisation of gravity and UV completion of GR
 - ▶ background independent quantisation, renormalisability, path integral
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**THANKS FOR
YOUR ATTENTION!**

